

# Fractional Cauchy problems on $\mathbb{S}_1^2$ and coordinates changed random fields

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## Abstract

We present some results about coordinates changed random fields and characterize the corresponding angular power spectrum. Thus, we provide a probabilistic interpretation of the anisotropies of cosmic microwave background (CMB) radiation in the theory of Big Bang.

We first introduce the fractional Cauchy problem on compact manifolds  $\mathcal{M}$  and therefore specialize such result on the unit sphere  $\mathcal{M} = \mathbb{S}_1^2$ . The unique strong solution to the fractional Cauchy problem is the transition law of a time changed Brownian motion on  $\mathbb{S}_1^2$  started at  $x \in \mathbb{S}_1^2$ , say  $\mathfrak{B}^x(\mathfrak{L}_t^\nu)$ ,  $t > 0$ , where  $\nu \in (0, 1)$  is the order of the time-fractional derivative and of the new clock  $\mathfrak{L}_t^\nu$  which is an inverse to a stable subordinator. We consider a random field on the sphere, say  $T(x)$ ,  $x \in \mathbb{S}_1^2$  and its spectral representation obtained by means of spherical harmonics. Thus, we define a coordinates changed random field on the sphere given by  $T(\mathfrak{B}^x(\mathfrak{L}_t^\nu)) = T(x + \mathfrak{B}^{x_N}(\mathfrak{L}_t^\nu))$  where  $x_N$  is the north pole. Our new random field is characterized by a new form of the angular power spectrum. In particular, it exhibits long range dependence structure due to the non-Markovian coordinates change. Moreover, we consider the subordinated Brownian motion  $\mathfrak{B}^x(\mathfrak{S}_t^\alpha)$  as coordinates change of  $T$ . In this case we obtain new random fields with short range dependence but interesting form of the new angular power spectrum. Interesting results are also obtained by considering  $T(x + \mathfrak{B}^{x_N}(S_t))$  where  $S_t$  is a subordinator with symbol  $\Psi$ .

CMB radiation is a thermal radiation filling the universe almost uniformly and can be affected by some anisotropies. In the analysis of the CMB radiation a key role is played by the angular power spectrum of the observed data, that is a realization of a random field on the sphere.

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